

# Supersymmetry and Extra Dimensions:

## Example Sheet 4

*Corrections and suggestions should be emailed to [f.quevedo@damtp.cam.ac.uk](mailto:f.quevedo@damtp.cam.ac.uk)*

**Exercise 4.1:** Consider the Schrödinger equation for a particle moving in two dimensions  $x$  and  $y$ . The second dimension is a circle of radius  $r$ . The potential corresponds to a square well ( $V(x) = 0$  for  $x \in (0, a)$  and  $V = \infty$  otherwise). Derive the energy levels for the two-dimensional Schrödinger equation and compare the result with the standard one-dimensional situation in the limit  $r \ll a$ .

**Exercise 4.2:** Consider the following Lagrangian

$$\mathcal{S} = \int d^4x \left( \frac{1}{g^2} H_{\mu\nu\rho} H^{\mu\nu\rho} + a \epsilon^{\mu\nu\rho\sigma} \partial_\mu H_{\nu\rho\sigma} \right).$$

Solve the equation of motion for the Lagrange multiplier  $a$  to obtain an action for a propagating massless Kalb-Ramond field  $B_{\mu\nu}$ . Alternatively, solve the equation of motion for the field  $H_{\nu\rho\sigma}$ , to obtain an action for the propagating axion field  $a$ . What happens to the coupling  $g$  under this transformation? Generalise your result for arbitrary dimensions and ranks of the tensors.

**Exercise 4.3:** Consider a massive antisymmetric tensor of rank  $q$  in  $D$  dimensions. Write up its Lagrangian up to second derivatives. Describe a general Lagrangian that can reproduce the original Lagrangian and its dual. Determine the degrees of freedom of the original and dual tensors. Interpret this dualisation in terms of a functional Fourier transform. Can this also be used in the massless case?

**Exercise 4.4** On spacetimes with Lorentzian signature show that only in dimensions  $D = 4k + 2$  there can be self-dual antisymmetric tensors. How many degrees of freedom do they have? What kind of  $p$ -branes they couple to? Explain the difference, if any, with Euclidean spaces.

**Exercise 4.5:** Show that the Kaluza-Klein dimensional reduction from  $D = 5$  to  $D = 4$  follows from a pure gravitational theory in five-dimensions, using  ${}^{(5)}R = {}^{(4)}R - 2e^{-\sigma}\nabla^2 e^\sigma - \frac{1}{4}e^{2\sigma}F_{\mu\nu}F^{\mu\nu}$  where  $G'_{55} = e^{2\sigma}$ . Relate the gauge transformation to the  $U(1)$  isometry of the compact space.

**Exercise 4.6** Demonstrate that the volume of a  $N - 1$  sphere of radius  $r$  is

$$V_{N-1} = \frac{2\pi^{N/2}}{\Gamma(N/2)} r^{N-1} \quad (1)$$

*Hint: It may help to consider the integral  $I_N = \int d^N x e^{-\rho^2}$  with  $\rho^2 = \sum_{i=1}^N x_i^2$ . Use this result to derive an expression for the electric (and gravitational) potential in  $D$  dimensions. Show that the potential due to a point particle in five dimensions reduces to the 4-dimensional potential at distances much larger than the size of the fifth dimension.*

**Exercise 4.7:** Consider a five dimensional gravity theory with a negative cosmological constant  $\Lambda < 0$ , compactified on an interval  $(0, \pi)$ . Each end of the interval corresponds to a '3-brane' which we choose to have tension  $\pm\Lambda/k$  respectively. Here  $k$  is a common scale to be determined later in terms of the fundamental scale in 5D  $M_*$  and  $\Lambda$ . Verify that the warped metric

$$ds^2 = e^{-2A(\theta)} \eta_{\mu\nu} dx^\mu dx^\nu - r^2 d\theta^2$$

satisfies Einstein's equations. Here  $W = e^{-2A(\theta)}$  is the warp factor and  $r$  is a constant measuring the size of the interval. You can use that Einstein's equations reduce to

$$\frac{6A'^2}{r^2} = -\frac{\Lambda}{2M_*^3}, \quad \frac{3A''}{r^2} = \frac{\Lambda}{2M_*^3 kr} [\delta(\theta - \pi) - \delta(\theta)].$$

Solve for  $A(\theta)$  and use the warp factor to show that the effective 4D Planck scale is now

$$M_{\text{pl}}^2 = M_*^3 r \int_{-\pi}^{\pi} d\theta e^{-2A} = \frac{M_*^3}{k} (1 - e^{-2kr}).$$

Find the value of the constant  $k$ . Consider the Higgs Lagrangian on the brane at  $\theta = \pi$ , bring it into canonical form and show that the mass is proportional to the factor  $e^{-k\pi r}$ . How large can  $r$  be in order to reproduce the electroweak scale from the Planck scale? Does this solve the hierarchy problem? How does the Planck scale differ from the 5D scale  $M_*$ ?

**Exercise 4.8:** Imagine that it were possible to have particles with all possible spins up to  $j = 3$ . What would the maximum dimensionality of spacetime be consistent with supersymmetry?

**Exercise 4.9:** Starting with the field contents of IIA and IIB supergravities in  $D = 10$  perform the dimensional reduction to  $D = 9$  and count the number of degrees of freedom for each multiplet. Is the spectrum chiral? Perform directly the reduction from  $D = 11$  to  $D = 9$  and compare. Perform dimensional reduction of IIB supergravity in  $D = 10$  all the way to  $D = 4$  and compare the number of degrees of freedom.

**Exercise 4.10:** Consider  $\mathcal{N} = 1$  supergravity with three chiral superfields  $S$ ,  $T$ , and  $C$ . In Planck units, the Kähler potential and superpotential are given by

$$\begin{aligned} K &= -\log(S + S^*) - 3 \log(T + T^* - CC^*) \\ W &= C^3 + a e^{-\alpha S} + b, \end{aligned}$$

where  $a, b$  are arbitrary complex numbers and  $\alpha > 0$ . Compute the scalar potential. Find the auxiliary field for  $S, T, C$  and verify that supersymmetry is broken. Assuming that  $C$  denotes a matter field with vanishing vev, find a minimum of the potential. Are there flat directions? A typical Kähler potential derived from string compactifications takes the form

$$K = -3 \log \Gamma(\tau_i) \tag{2}$$

where  $\Gamma$  is a homogeneous function of degree one of moduli fields  $\tau_i$ . Using the homogeneity equations  $\tau_i \Gamma_i = \Gamma$  and  $\tau_i \Gamma_{ij} = 0$  (where  $\Gamma_i = \partial \Gamma / \partial \tau_i$ , etc.) show that

$$\tau_i K_{ij} = 3 \Gamma_j / \Gamma, \quad \Gamma_i K_{ij}^{-1} \Gamma_j / \Gamma^2 = 1/3$$

and deduce from this that if the superpotential does not depend on the  $\tau_i$  fields then the corresponding contribution to the  $N = 1$  supergravity scalar potential  $V$  vanishes.